

BNL spin discussion
09/09/03

A_{TT} in NLO - a status report

Marco Stratmann

Univ. of Regensburg



work done in collaboration with [A. Mukherjee](#)
and [W. Vogelsang](#)

Accessing δq at RHIC

chirality has to be flipped *twice* to access δq

↙ ↘
double-spin single-spin
asymmetries

A_{TT} A_N

A_{TT} :

✓ clean, involves only δq as “unknowns”

✗ expected to be small → exp. challenge

“selection rule” $A_{TT} \ll A_{LL}$

Artru, Mekhfi; Ji; Jaffe, Saito

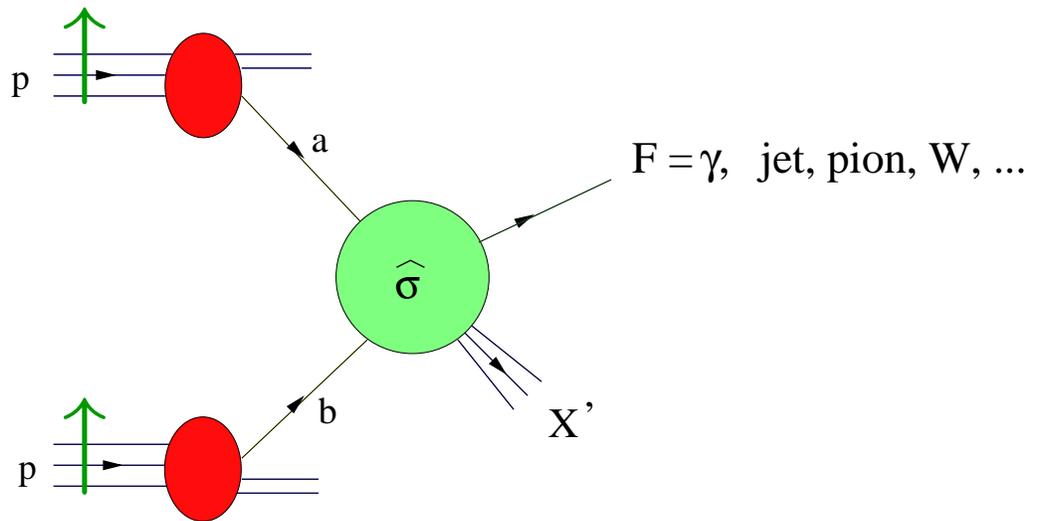
A_N :

✗ requires *other unknown* chiral-odd fct. → involved

Pandora’s box: Collins fct., Sivers fct, ...

✓ $pp^\uparrow \rightarrow \pi X$ sizable (E704, STAR)

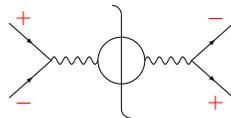
A_{TT} at RHIC



$$A_{TT} = \frac{d\sigma^{p^\uparrow p^\uparrow} - d\sigma^{p^\uparrow p^\downarrow}}{d\sigma^{p^\uparrow p^\uparrow} + d\sigma^{p^\uparrow p^\downarrow}} \propto \sum_{a,b} \delta f_a \otimes \delta f_b \otimes \underbrace{\frac{\delta \hat{\sigma}^{ab \rightarrow FX'}}{p\text{QCD.}}}_{\text{pQCD.}}$$

on the menu:

- **Drell-Yan process**



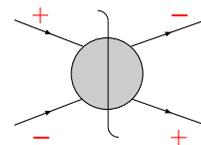
Ralston, Soper; Ji

Cortes, Pire, Ralston; Artru, Mekhfi; Jaffe, Ji

- most suited process: no gluons in LO
- NLO study: meas. suffers from limited μ^\pm acceptance nevertheless, appears feasible

Martin, Schäfer, MS, Vogelsang

- **high- p_T prompt photons, hadrons, jets**

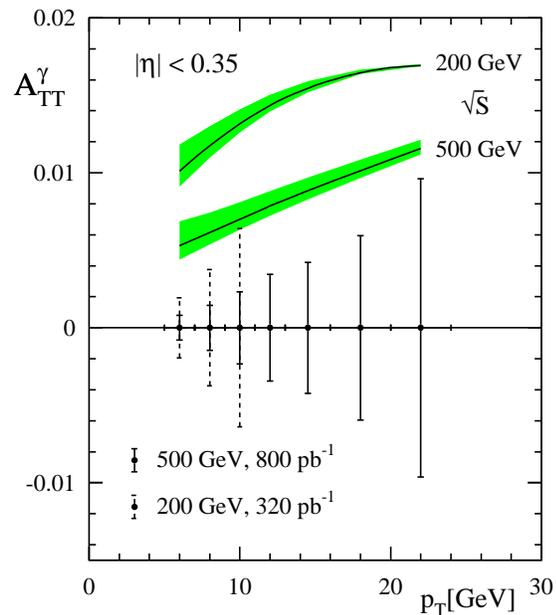
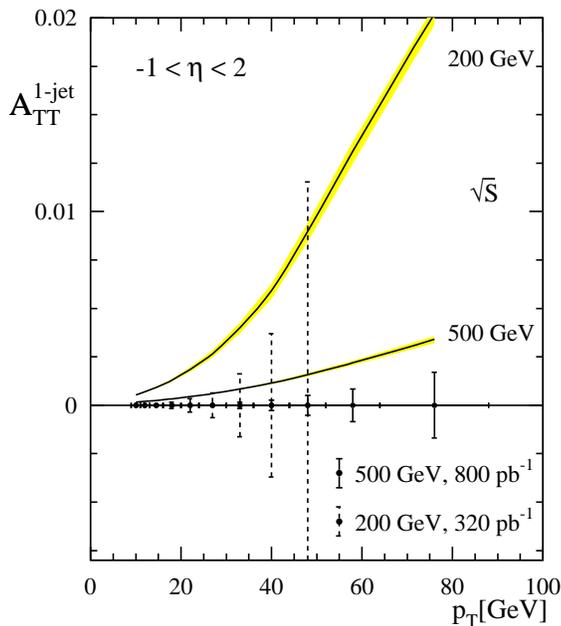


- A_{TT} small due to absence of $g^\uparrow g^\uparrow$ and $q^\uparrow g^\uparrow$ processes
- sizable rates \rightarrow statistics sufficient even if A_{TT} small

LO results assuming $\delta q(x, \mu_0) = \frac{1}{2} [q(x, \mu_0) + \Delta q(x, \mu_0)]$
 Soffer, MS, Vogelsang

$p^\uparrow p^\uparrow \rightarrow \text{jet } X$
 STAR

$p^\uparrow p^\uparrow \rightarrow \gamma X$
 PHENIX



however, NLO QCD corrections are in general a must:

(scale dependence, ...)

→ Barbara's talk

further motivation for NLO: "technical challenge"

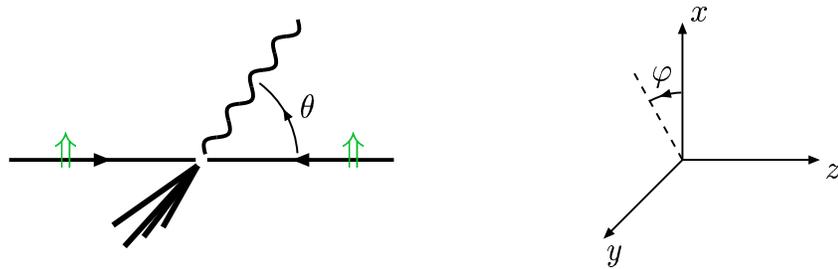
- general paucity of NLO calculations with transverse spin (until recently: NLO only for incl. DY and evol. kernels)
- provide and apply a feasible technique

Why transverse spin is more complicated to handle:

long. polarization: spin *aligned* with momentum ✓

trans. polarization: spin = *extra spatial direction*

↔ non-trivial azimuthal dep.



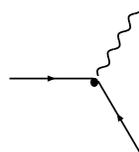
always of the form $\frac{d^3\delta\sigma}{dp_T d\eta d\Phi} \equiv \cos(2\Phi) \left\langle \frac{d^2\delta\sigma}{dp_T d\eta} \right\rangle$

→ ϕ integration not appropriate

but standard NLO techniques rely on integrations over

full azimuthal phase space *plus* use of particular

reference frame



Gottfried, Jackson

⇒ difficult to deal with azimuthal angle

(in particular, in $d = 4 - 2\epsilon$ dimensions)

recent progress:

NLO corrections to $p^\uparrow p^\uparrow \rightarrow \gamma X$

A. Mukherjee, MS, W. Vogelsang

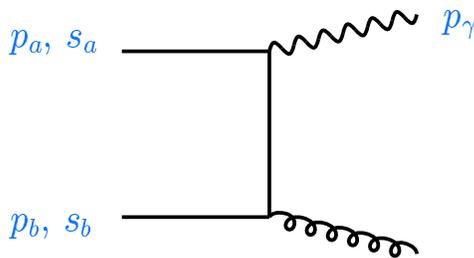
key point: ϕ -dep. always results from **covariant** expression

$$\mathcal{F}(p_\gamma, s_a, s_b) = \frac{s}{tu} \left[2 (p_\gamma \cdot s_a) (p_\gamma \cdot s_b) + \frac{tu}{s} (s_a \cdot s_b) \right]$$

$$= \cos(2\phi) \text{ in hadronic c.m.s.}$$

\Rightarrow use \mathcal{F} to project out ϕ covariantly

LO example: $q\bar{q} \rightarrow \gamma g$



$$p_a \cdot s_a = p_b \cdot s_b = 0$$

$$s_a^2 = s_b^2 = -1$$

matrix element [use $u(p_a, s_a) \bar{u}(p_a, s_a) = \frac{1}{2} \not{p}_a [1 + \gamma_5 \not{s}_a], \dots$]

$$\delta|M|^2 = (ee_{qg})^2 \frac{4C_F}{N_C} \frac{s}{tu} \left[2 (p_\gamma \cdot s_a) (p_\gamma \cdot s_b) + \frac{tu}{s} (s_a \cdot s_b) \right]$$

project with \mathcal{F} :

$$\frac{1}{\pi} \int d\Omega_\gamma \mathcal{F}(p_\gamma, s_a, s_b) \delta|M|^2 = (ee_{qg})^2 \frac{4C_F}{N_C} = \langle \delta|M|^2 \rangle \checkmark$$

terms involving $p_\gamma \cdot s_a, p_\gamma \cdot s_b$ can be integrated “covariantly”

easily generalized to NLO calculation in d dimensions:

(1) multiply any $\delta|M|^2$ with $\mathcal{F}(p_\gamma, s_a, s_b)$

(2) integrate all resulting scalar products with $s_{a,b}$

for example:

[up to $\mathcal{O}(\epsilon)$]

$$\int d\Omega_\gamma (p_\gamma \cdot s_a)^2 (p_\gamma \cdot s_b)^2 = \int d\Omega_\gamma \frac{t^2 u^2}{8s^2} [2(s_a \cdot s_b)^2 + s_a^2 s_b^2]$$

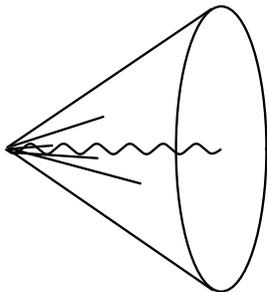
(3) arrive at a structure similar to an unpol. $|M|^2$

(4) employ *standard techniques* for phase space integr.

(5) restore ϕ dependence afterwards

remarks:

- cancellation of divergencies proceeds as usual
→ Barbara's talk
- at colliders, impose "isolation cut" on photon:



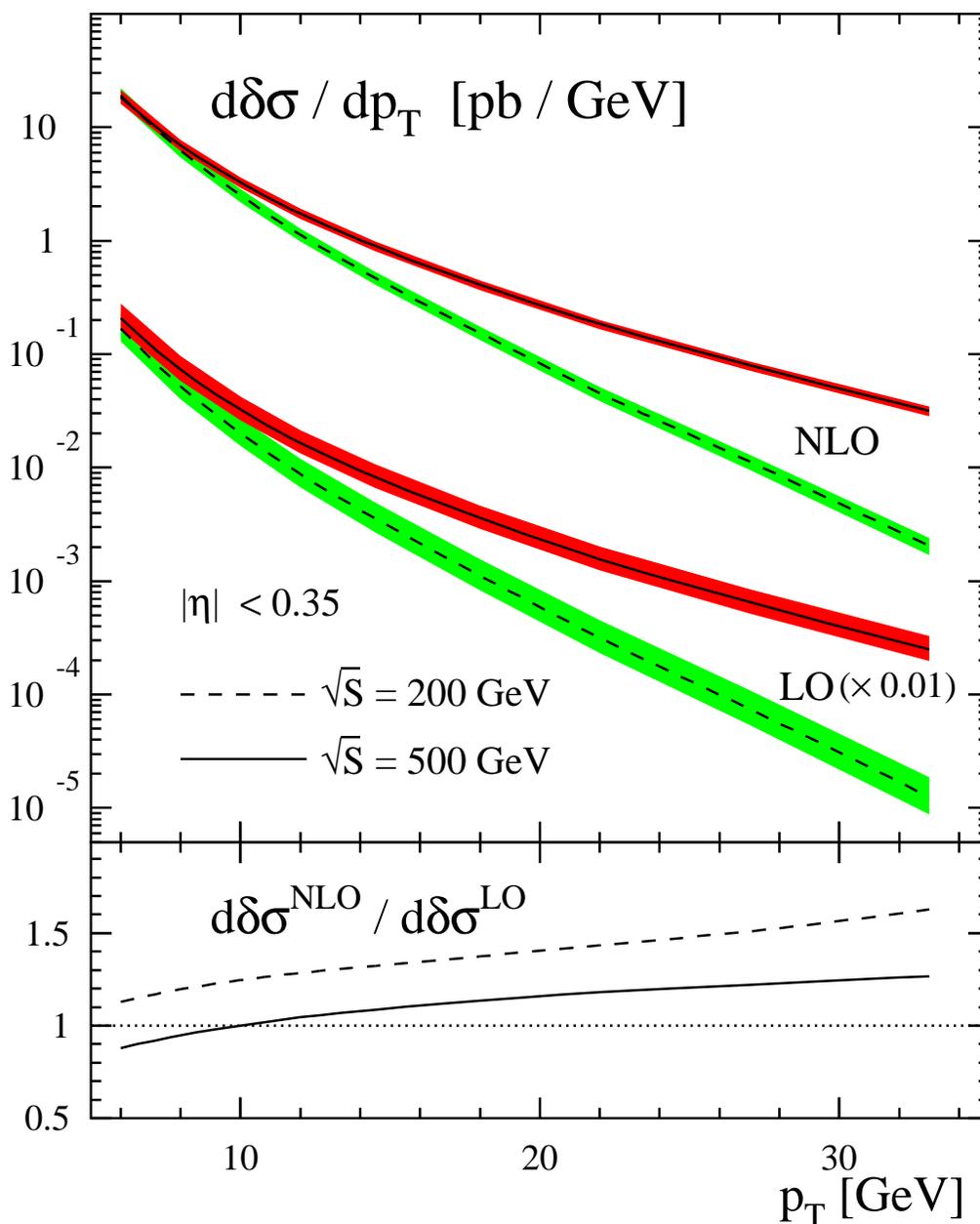
typically, demand $E_{\text{had}} \leq \epsilon E_\gamma$

in $\sqrt{(\Delta\phi)^2 + (\Delta\eta)^2} \leq R$

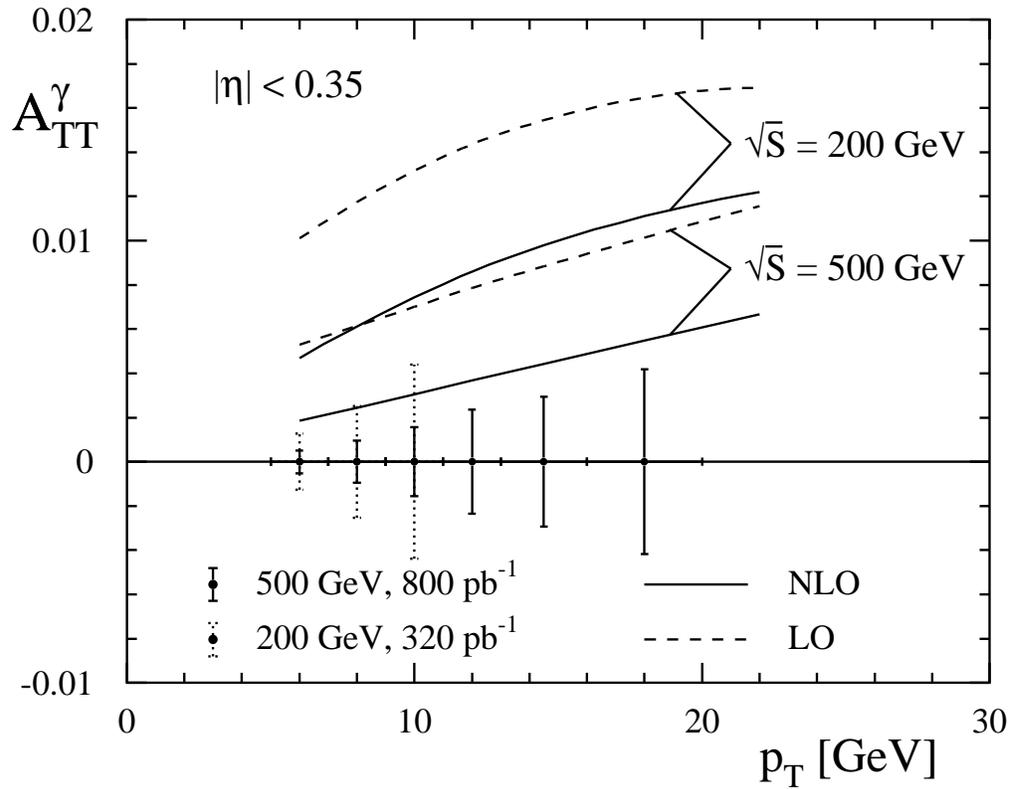
- for isolation with $\epsilon \propto (r/R)^2$ one can eliminate the fragmentation component altogether Frixione

results for $p^\uparrow p^\uparrow \rightarrow \gamma X$:

- improved scale dependence
- reasonable “K-factors”



results for A_{TT}^γ :



A. Mukherjee, MS, W. Vogelsang

work in progress:

- half-way through with $p^\uparrow p^\uparrow \rightarrow \pi X$
- future: $p^\uparrow p^\uparrow \rightarrow \text{jet} X \dots$